

# No muon excess in EAS at 100 PeV: results of EAS-MSU experiment

Mikhail Kuznetsov

In collaboration with

S. Troitsky, G. Rubtsov, I. Karpikov (INR RAS)

Yu. Fomin, N. Kalmykov, G. Kulikov and V. Sulakov (SINP MSU)

*Astropart.Phys.* 92, 1 (2017)



**INR**

Institute for Nuclear Research  
of the Russian Academy of Sciences

ICRC-2017, Busan, July 2017

- Motivation: study of muons in extensive air showers (EAS)
- The EAS-MSU experiment
- The method of muon excess description
- Results & Conclusions

## Motivation: study of muons in EAS

- Muon content of EAS as a tool for study the hadronic interactions at high energies.
- At highest energies ( $E_{CR} > 10^{19}$  eV ) the presence of “muon excess” is well established.
- The cosmic-ray energy  $\sim 10^{19}$  eV correspond to  $\sim 100$  TeV LHC energy.

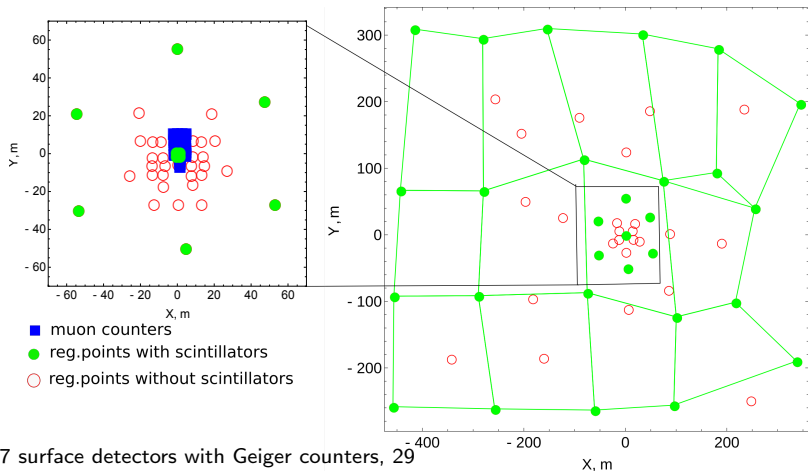
The “muon excess” is the overabundance of muons in a real EAS comparing to high-energy hadronic models predictions.

### Published results for EAS muon content in various parameter regions

| Experiment | $X$ ,<br>$\text{g}/\text{cm}^2$ | $E$ , eV            | $E_\mu$ ,<br>GeV | $r/r_0$      | $\theta$        | $\mu$ excess |
|------------|---------------------------------|---------------------|------------------|--------------|-----------------|--------------|
| HiRes-MIA  | 860                             | $10^{17} - 10^{18}$ | $\gtrsim 0.85$   | $\gtrsim 10$ | N/A             | yes          |
| PAO        | 880                             | $\gtrsim 10^{19}$   | $\gtrsim 1$      | $\gtrsim 10$ | $70^\circ$      | yes          |
| Yakutsk    | 1020                            | $\gtrsim 10^{19}$   | $\gtrsim 1$      | $\gtrsim 10$ | $45^\circ$      | yes          |
| IceTop     | 680                             | $10^{15} - 10^{17}$ | $\gtrsim 0.2$    | $\gtrsim 3$  | $13^\circ$ mean | no           |

For TA muon results see R. Takeishi poster on 18.07

# EAS-MSU experiment



- 77 surface detectors with Geiger counters, 29 scintillators for timing
- $\sim 0.5 \text{ km}^2$  area
- Underground muon detector ( $E_\mu \geq 10 \text{ GeV}$ )
- $\sim 1500$  days of operation (1982-1990)
- Total  $\sim 10^6$  EASes detected

Effective for EAS with

$$10^{15.5} \lesssim E_{\text{primary}} \lesssim 10^{17.75} \text{ eV}$$

## Full Monte-Carlo simulation

Yu. Fomin et al., JINST 11, T08005 (2016)

CORSIKA (p + Fe primaries) + EGS4 + Fluka + QGSJetII

$$10^{15.5} \leq E_{\text{primary}} \leq 10^{17.75} \text{ eV}, \quad \frac{dN}{dE} \sim E^{-3.1}$$

avg. of PDG 2014 spectra



C++ code simulating the facility



Reconstruction with the same code as for data

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Reconstruction with the same code as for data

## SD reconstruction parameters

- $N_e$  — number of charged particles
- $S$  — shower age parameter
- $\theta, \phi$  — shower zenith and azimuthal angle
- $R$  — shower core distance from the array center

All particle lateral distribution function (LDF)

$$\Rightarrow \rho(S, r) \sim N_e \cdot \left(\frac{r}{r_0}\right)^{(S+\alpha(r/r_0)-2)} \times \left(\frac{r}{r_0} + 1\right)^{(S+\alpha(r/r_0)-4.5)}$$

$r_0 \simeq 80\text{m}$  — is the Moliere radius



# Monte-Carlo & reconstruction II

## Selection cuts

- $N_e > 2 \cdot 10^7$  ( $E_{MC} \gtrsim 10^{16.5}$  eV for p-primaries)
- $R < 240\text{m}$
- $\theta < 30^\circ$
- $0.3 < S < 1.8$

## $1\sigma$ accuracy of the reconstruction for p/Fe mix

- $\Delta\psi < 1.1^\circ$  — arrival direction
- $\Delta R < 5.7\text{ m}$  — axis position
- $\Delta N_e/N_e < 0.17$  — number of charged particles
- The efficiency is  $\gtrsim 95\%$

Total exposure is  $7.7 \times 10^6$  km s sr,

809 events selected

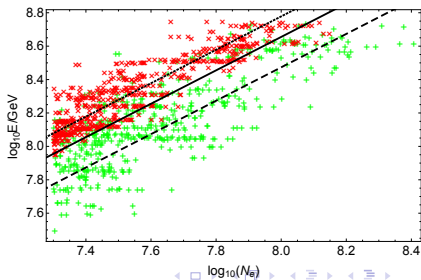
## Primary energy reconstruction

For best fit p/Fe mix:

$$\log(E/\text{GeV}) = 0.65 + \log_{10} N_e$$

$\Downarrow$

$$\Delta E/E < 0.41 \quad (1\sigma)$$

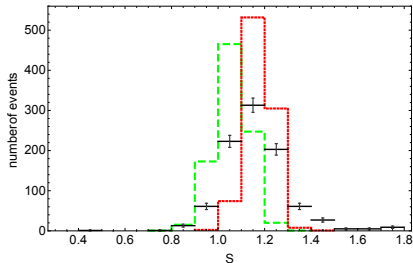


# Monte-Carlo: SD chemical composition

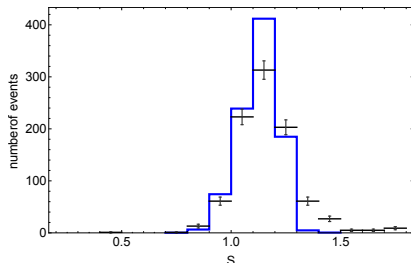
Chemical composition from the fit of SD data over LDF  $S$ -parameter.

$$\rho(S, r) \sim N_e \cdot \left(\frac{r}{r_0}\right)^{(S+\alpha(r/r_0)-2)} \times \left(\frac{r}{r_0} + 1\right)^{(S+\alpha(r/r_0)-4.5)}$$

Data vs. MC p and MC Fe



Data fitted with MC p/Fe mix



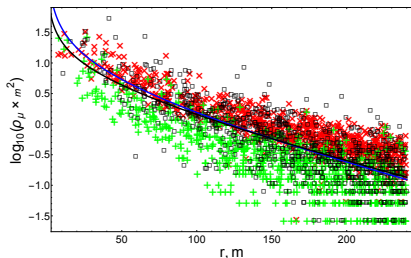
$$\frac{N_p}{N_p + N_{Fe}} = 43\%, \quad \frac{N_{Fe}}{N_p + N_{Fe}} = 57\%$$

Cf. the results of other experiments. KASCADE-Grande: 41% p , 59% Fe .  
Tunka-133: 49% p , 51% Fe .



## Muon measurement & MC

- Muon detector locates in the center of array at 40 m w.e. depth
- The muon energy threshold is  $E_\mu \geq 10$  GeV
- Consists of Geiger counters, total area is  $\sim 36\text{m}^2$

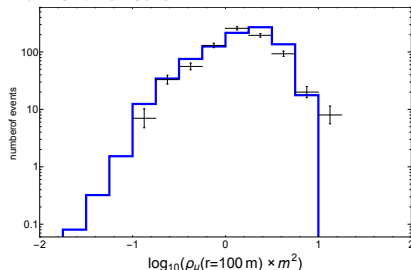


### Results for $\rho_\mu$ at 100 m from shower core

Use  $\mu$ LDF to calculate  $\rho_\mu(R = 100\text{m})$ :

$$\rho_\mu(r) = N_\mu \left( \frac{r}{r_0} \right)^{-a_\mu} \exp(-r/r_0), \quad a_\mu = 0.7$$

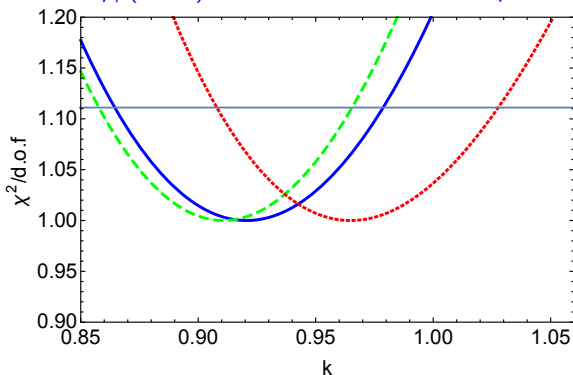
Data agree well with Monte-Carlo



## Test for muon excess

Make MC sets of EASes with artificially multiplied or reduced number of muons:  $\widetilde{N}_\mu = k \cdot N_\mu$

Results of  $\rho_\mu(100\text{m})$  fit for various chemical composition



EAS-MSU SD:  $k = 0.92 \pm 0.06$ ,  $k > 1$  excluded at 92 % C.L.

KASCADE-Grande:  $k = 0.91 \pm 0.05$ ,  $k > 1$  excluded at 95 % C.L.

Tunka-133:  $k = 0.96 \pm 0.06$ ,  $k > 1$  excluded at 67 % C.L.

**No muon excess**

## Comparison with results of other experiments

| Experiment     | $X$ ,<br>$\text{g}/\text{cm}^2$ | $E$ , eV                              | $E_\mu$ ,<br>GeV               | $r/R_0$                        | $\theta$                     | $\mu$ excess |
|----------------|---------------------------------|---------------------------------------|--------------------------------|--------------------------------|------------------------------|--------------|
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| <b>EAS-MSU</b> | <b>990</b>                      | <b><math>10^{17} - 10^{18}</math></b> | <b><math>\gtrsim 10</math></b> | <b><math>\lesssim 3</math></b> | <b><math>30^\circ</math></b> | <b>no</b>    |

The results are not contradictory but complementary.

Test  $\mu$ -excess for other interaction models (simplified Monte-Carlo)

| Model        | mean Fe fraction<br>from $S$ , % | $k$             |
|--------------|----------------------------------|-----------------|
| QGSJET-II-04 | 57                               | $0.92 \pm 0.06$ |
| EPOS-LHC     | 42                               | $0.96 \pm 0.06$ |
| SIBYLL 2.1   | 76                               | $1.00 \pm 0.07$ |
| QGSJET-01    | 58                               | $0.95 \pm 0.06$ |

## Conclusions

- EAS-MSU archival data was studied with full Monte-Carlo.
- There is no “muon excess” at  $10^{17} \lesssim E_{CR} \lesssim 10^{18}$  eV,  $E_{\mu} \geq 10$  GeV and in the central part of the shower.

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## Addition

The search for  $\sim 100$  PeV photons was also performed with EAS-MSU data, the stringent flux upper-limits was set.

See my poster on 18.07

# Thank you!

## Simplified Monte-Carlo for EPOS, QGSJet-01 and Sibyll

- Simulation of 800 p and 800 Fe CORSIKA events with  $E = 10^{17}$  and  $E = 4 \cdot 10^{17}$  eV with  $\varepsilon = 10^{-5}$  thinning.
- For thinning the maximal e/m and hadronic weights are equal to 4000 and 40 respectively.
- The equivalence between thinned and non-thinned showers was justified.
- The age parameter  $S$  was calculated by averaging the particle density over concentric rings around the shower axis.
- The muon number  $N_\mu$  was calculated by counting CORSIKA muons.

## The EAS-MSU statics is not enough to study:

- Dependence of  $\mu$ LDF upon variation of zenith angle
- Dependence of the result upon spectral features (like “second knee”).
- Chemical composition beyond p/Fe mix approximation